# Some relations between rectifying and normal curves in Minkowski 3-space 

Milica Grbović ${ }^{1}$ and Emilija Nešović ${ }^{1, *}$<br>${ }^{1}$ Department of Mathematics and Informatics, University of Kragujevac, Radoja<br>Domanovića 12, SRB-34 000 Kragujevac, Serbia

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#### Abstract

In this paper, we obtain explicit parameter equations of spacelike rectifying curves in $E_{1}^{3}$ whose projection onto spacelike, timelike and lightlike plane of $E_{1}^{3}$ is a normal curve. We also obtain explicit parameter equations of spacelike normal curves in $E_{1}^{3}$ whose projection onto lightlike plane of $E_{1}^{3}$, with respect to a chosen screen distribution, is a rectifying W-curve. AMS subject classifications: 53C50, 53C40


Key words: Minkowski 3-space, rectifying curve, normal curve, curvature

## 1. Introduction

In the Euclidean space $\mathbb{E}^{3}$ there exist three classes of curves, so-called rectifying, normal and osculating curves satisfying Cesaro's fixed point condition ([10]) meaning that rectifying, normal and osculating planes of such curves always contain a particular point. If all normal or osculating planes of a curve in $\mathbb{E}^{3}$ pass through a particular point, then the curve is spherical or planar, respectively. It is also known that if all rectifying planes of a non-planar curve in $\mathbb{E}^{3}$ pass through a particular point, then the ratio of torsion and curvature of such curve is a non-constant linear function ([1]). Some characterizations of rectifying curves in Minkowski 3-space $E_{1}^{3}$ are given in [7]. In particular, there exists a simple relationship between rectifying curves and Darboux vectors (centrodes), which play some important roles in mechanics, kinematics as well as in differential geometry in defining the curves of constant precession ([2]). Normal curves in Minkowski 3-space are characterized in [5,6]. Spacelike and timelike normal curves in $E_{1}^{3}$ always lie in some quadric and null normal curves in $E_{1}^{3}$ are the null straight lines.

It is a quite interesting problem to obtain explicit parameter equations of rectifying and normal curves in Minkowski 3-space. In order to obtain such equations, it is natural to impose some extra condition on the corresponding curve. In this paper, we obtain explicit parameter equation of spacelike rectifying curve in $E_{1}^{3}$ assuming that orthogonal projection of such curve onto spacelike or timelike plane of $E_{1}^{3}$ is a normal curve. We prove that the straight lines are the only rectifying curves in $E_{1}^{3}$ whose projection onto lightlike plane with respect to a chosen screen distribution is

[^0]a normal curve. Also, we find explicit parameter equation of non-planar spacelike normal curve in $E_{1}^{3}$ assuming that the projection of such curve onto lightlike plane of $E_{1}^{3}$ (with respect to a chosen screen distribution) is a rectifying W-curve.

## 2. Preliminaries

The Minkowski 3 -space $\mathbb{E}_{1}^{3}$ is the Euclidean 3 -space $\mathbb{E}^{3}$ equipped with the indefinite flat metric given by

$$
g=-d x_{1}^{2}+d x_{2}^{2}+d x_{3}^{2}
$$

where $\left(x_{1}, x_{2}, x_{3}\right)$ is a rectangular coordinate system of $\mathbb{E}_{1}^{3}$. Recall that an arbitrary vector $v \in \mathbb{E}_{1}^{3}$ can be spacelike if $g(v, v)>0$ or $v=0$, timelike if $g(v, v)<0$ and null (lightlike) if $g(v, v)=0$ and $v \neq 0([3,9])$. The norm of a vector $v$ is given by $\|v\|=\sqrt{|g(v, v)|}$ and two vectors $v$ and $w$ are said to be orthogonal, if $g(v, w)=0$. An arbitrary curve $\alpha(s)$ in $\mathbb{E}_{1}^{3}$, can locally be spacelike, timelike or null (lightlike), if all its velocity vectors $\alpha^{\prime}(s)$ are spacelike, timelike or null, respectively ([9]). A spacelike or timelike curve $\alpha(s)$ has unit speed, if $g\left(\alpha^{\prime}(s), \alpha^{\prime}(s)\right)= \pm 1$. Arbitrary curve in $E_{1}^{3}$ is called $W$-curve, if all its curvature functions are constant.

Let $\{T, N, B\}$ be the moving Frenet frame along a curve $\alpha$ in $\mathbb{E}_{1}^{3}$, consisting of the tangent, principal normal and binormal vector field, respectively. If $\alpha$ is a non-null curve with non-null principal normal $N$, the Frenet equations read ([8])

$$
\left[\begin{array}{l}
T^{\prime}  \tag{1}\\
N^{\prime} \\
B^{\prime}
\end{array}\right]=\left[\begin{array}{ccc}
0 & \epsilon_{2} \kappa_{1} & 0 \\
-\epsilon_{1} \kappa_{1} & 0 & -\epsilon_{1} \epsilon_{2} \kappa_{2} \\
0 & -\epsilon_{2} \kappa_{2} & 0
\end{array}\right]\left[\begin{array}{l}
T \\
N \\
B
\end{array}\right]
$$

where $\kappa_{1}(s), \kappa_{2}(s)$ are the first and second curvature of the curve, $\epsilon_{1}=g(T, T)= \pm 1$, $\epsilon_{2}=g(N, N)= \pm 1$ and $g(B, B)=-\epsilon_{1} \epsilon_{2}$, respectively.

If $\alpha$ is a spacelike curve with null principal normal $N$, it is called pseudo null curve and the Frenet equations have the form ([11])

$$
\left[\begin{array}{l}
T^{\prime}  \tag{2}\\
N^{\prime} \\
B^{\prime}
\end{array}\right]=\left[\begin{array}{ccc}
0 & \kappa_{1} & 0 \\
0 & \kappa_{2} & 0 \\
-\kappa_{1} & 0 & -\kappa_{2}
\end{array}\right]\left[\begin{array}{l}
T \\
N \\
B
\end{array}\right]
$$

where the first curvature $\kappa_{1}(s)=0$ if $\alpha$ is a straight line, or $\kappa_{1}(s)=1$ in all other cases. In this case, the following conditions hold

$$
g(T, T)=1, g(N, N)=g(B, B)=0, g(T, N)=g(T, B)=0, g(N, B)=1
$$

Recall that the normal and rectifying curves in $E_{1}^{3}$ are defined in [6] and [7] as the curves whose the position vector $\alpha$ (with respect to some chosen origin) always lies in its normal plane $T^{\perp}$ and rectifying plane $N^{\perp}$, respectively. Therefore, the position vector of normal or rectifying curve $\alpha$ in $E_{1}^{3}$ satisfy the equations $g(\alpha, T)=0$, $g(\alpha, N)=0$, respectively.

Let $\Omega$ be the lightlike plane of $E_{1}^{3}$. Denote by $T \Omega$ the tangent bundle subspace of $\Omega$ in $E_{1}^{3}$. Since $\left.g\right|_{\Omega}$ is degenerate on $\Omega$, the radical or the null space of $T_{p} \Omega$ at each point $p \in \Omega$, is a subspace $\operatorname{Rad} T_{p} \Omega$ defined by ([3])

$$
\operatorname{Rad} T_{p} \Omega=\left\{Y \in T_{p} \Omega \mid g_{p}(X, Y)=0, \forall X \in T_{p} \Omega\right\}
$$

For the radical space there holds

$$
\operatorname{Rad} T_{p} \Omega=T_{p} \Omega \cap T_{p} \Omega^{\perp}
$$

Moreover, since $\Omega$ is lightlike plane in $E_{1}^{3}$, it follows that $\operatorname{dim}\left(T_{p} \Omega^{\perp}\right)=1$ and therefore $\operatorname{dim}\left(\operatorname{Rad} T_{p} \Omega\right)=1$ and $\operatorname{Rad} T_{p} \Omega=T_{p} \Omega^{\perp}$. The vector bundle $\operatorname{Rad} T \Omega$ is called the radical (null) distribution of $\Omega$.

Denote by $S(T \Omega)$ a complementary vector bundle of $\operatorname{Rad} T \Omega$ in $T \Omega$. This means that

$$
T \Omega=\operatorname{Rad} T \Omega \oplus S(T \Omega)
$$

Vector bundle $S(T \Omega)$ is called a screen distribution on $\Omega$. In particular, for a given screen distribution $S(T \Omega)$ there exists a unique complementary vector bundle $\operatorname{ltr}(T \Omega)$ to $T \Omega$ in $\left.T E_{1}^{3}\right|_{\Omega}$. The vector bundle $\operatorname{ltr}(T \Omega)$ is called the lightlike transversal vector bundle of $\Omega$. Consequently, the tangent bundle $T E_{1}^{3}$ splits into the following three non-intersecting complementary (but not orthogonal) vector bundles ([3]):

$$
\left.T E_{1}^{3}\right|_{\Omega}=\operatorname{Rad} T \Omega \oplus S(T \Omega) \oplus \operatorname{ltr}(T \Omega)
$$

Let us consider arbitrary curve $\alpha$ lying fully in $E_{1}^{3}$ (and not lying in its lightlike plane) and its projection $\beta=P^{*}(\alpha)$ onto lightlike plane $\Omega$ of $E_{1}^{3}$, where $P^{*}: E_{1}^{3} \rightarrow \Omega$ is the projection map. Then the projection map $P^{*}$ is not orthogonal and unique ([4]). In particular, the curve $\alpha(s)=\left(\alpha_{1}(s), \alpha_{2}(s), \alpha_{3}(s)\right)$ and the projected curve $\beta$ are related by

$$
\begin{equation*}
\alpha(s)=\beta(s)+g(\alpha(s), A) B \tag{3}
\end{equation*}
$$

where $A$ and $B$ are null vectors satisfying the conditions

$$
\begin{aligned}
\operatorname{Rad} T \Omega & =\operatorname{span}\{A\}, \quad \operatorname{ltr}(T \Omega)=\operatorname{span}\{B\} \\
g(A, B) & =1, \quad g(B, W)=0, \quad \forall W \in S(T \Omega)
\end{aligned}
$$

Next we show that by choosing different screen distributions on $\Omega$, different parameter equations of the projected curves $\beta=P^{*}(\alpha)$ can be obtained.

Up to isometries of $E_{1}^{3}$, assume that $\Omega$ has the equation $x_{1}=x_{2}$. Then the radical distribution $\operatorname{Rad} T \Omega$ is spanned by

$$
\begin{equation*}
A=(1,1,0) . \tag{4}
\end{equation*}
$$

Let $Q=(0,1,0)$ be locally defined nowhere zero section defined on the open subset $U \subset E_{1}^{3}$. Then $g(Q, A)=1, g(Q, Q)=1$ and therefore the lightlike transversal vector bundle is spanned by ([3])

$$
\begin{equation*}
B=\frac{1}{g(A, Q)}\left\{Q-\frac{g(Q, Q)}{2 g(A, Q)} A\right\}=\left(-\frac{1}{2}, \frac{1}{2}, 0\right) . \tag{5}
\end{equation*}
$$

It follows that the corresponding screen distribution $S(T \Omega)$ is spanned by

$$
\begin{equation*}
W=(0,0,1) \tag{6}
\end{equation*}
$$

Substituting (4) and (5) in (3), we obtain that the projected curve $\beta$ has parameter equation of the form

$$
\begin{equation*}
\beta(s)=\left(\frac{\alpha_{1}(s)+\alpha_{2}(s)}{2}, \frac{\alpha_{1}(s)+\alpha_{2}(s)}{2}, \alpha_{3}(s)\right) . \tag{7}
\end{equation*}
$$

Let $Q=(0,1,1)$ be the next locally defined nowhere zero section defined on the open subset $U \subset E_{1}^{3}$. Then $g(Q, A)=1, g(Q, Q)=2$, so relation (5) implies that the lightlike transversal vector bundle is spanned by

$$
\begin{equation*}
B=(-1,0,1) \tag{8}
\end{equation*}
$$

It follows that the corresponding screen distribution $S(T \Omega)$ is spanned by

$$
\begin{equation*}
V=(1,1,-1) \tag{9}
\end{equation*}
$$

Substituting (4) and (8) in (3), we find that the projected curve $\beta$ is given by

$$
\begin{equation*}
\beta(s)=\left(\alpha_{2}(s), \alpha_{2}(s), \alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)\right) \tag{10}
\end{equation*}
$$

Therefore, different choices of screen distributions on $\Omega$ provide different parameter equations of the projected curves (for more details, see [4]).

## 3. Some relations between spacelike rectifying curves and planar normal curves in $\mathbb{E}_{1}^{3}$

In this section we obtain some explicit parameter equations of spacelike rectifying curves in $E_{1}^{3}$. Recall that up to a parametrization, every spacelike rectifying curve with non-null principal normal in $E_{1}^{3}$ has one of the following parameter equations ([7]):

$$
\alpha(t)=\frac{a}{\cos t} x(t), \quad \beta(t)=\frac{b}{\sinh t} y(t), \quad \gamma(t)=\frac{c}{\cosh t} z(t),
$$

where $a, b, c \in R_{0}^{+}, x(t)$ is some unit speed spacelike curve lying in the pseudosphere $S_{1}^{2}(1), y(t)$ is some unit speed timelike curve lying in the pseudosphere $S_{1}^{2}(1)$ and $z(t)$ is some unit speed spacelike curve lying in the pseudohyperbolic space $H_{0}^{2}(1)$. In particular, rectifying curve $\alpha$ has spacelike rectifying plane, while the rectifying curves $\beta$ and $\gamma$ both have timelike rectifying planes and spacelike and timelike position vectors, respectively.

In the theorems which follow, we will determine explicit parameter equations of the curves $x(t), y(t)$ and $z(t)$ by imposing extra condition on the rectifying curve. Namely, we will assume that its orthogonal projection onto non-degenerate plane of $E_{1}^{3}$ is a normal curve.

Theorem 1. Let $\alpha$ be a spacelike rectifying curve in $E_{1}^{3}$ and $\beta$ the orthogonal projection of $\alpha$ onto spacelike plane of $E_{1}^{3}$.
(i) If $\alpha$ has non-null principal normal and $\beta$ is normal curve, then up to a parametrization $\alpha$ is given by

$$
\begin{equation*}
\alpha(u)=(c \cos u, \cos (c u), \sin (c u)), \tag{11}
\end{equation*}
$$

where $c \in R_{0}^{+}, c \neq 1$;
(ii) If $\alpha$ has null principal normal and $\beta$ is normal curve, then up to a parametrization $\alpha$ is given by

$$
\begin{equation*}
\alpha(u)=(\cos u, \cos u, \sin u) \tag{12}
\end{equation*}
$$

Proof. (i) Let $\beta$ be the orthogonal projection of $\alpha$ onto spacelike plane of $E_{1}^{3}$. Then $\alpha$ is given by

$$
\begin{equation*}
\alpha(s)=\beta(s)+\mu(s) \mathrm{v}, \tag{13}
\end{equation*}
$$

where $s$ is arclength parameter of $\alpha, \mu(s)$ is some non-constant differentiable function and $\mathrm{v}=(1,0,0)$ is a unit timelike vector.

Assume that $\alpha$ has non-null principal normal and that $\beta$ is a normal curve. Then the equation of $\beta$ reads

$$
\begin{equation*}
\beta(t)=(0, \cos t, \sin t) \tag{14}
\end{equation*}
$$

where $t$ is arclenght parameter of $\beta$. Differentiating relation (13) with respect to $s$ and using the condition $g\left(\alpha^{\prime}(s), \alpha^{\prime}(s)\right)=1$, we get

$$
\begin{equation*}
1+\mu^{\prime 2}(s)=t^{\prime 2}(s) \tag{15}
\end{equation*}
$$

Differentiating relation (14) with respect to $s$ we find

$$
\begin{equation*}
g\left(\beta(s), \beta^{\prime \prime}(s)\right)=-t^{2}(s) \tag{16}
\end{equation*}
$$

which together with the condition $g\left(\alpha(s), \alpha^{\prime \prime}(s)\right)=0$ yield

$$
\begin{equation*}
t^{\prime 2}(s)+\mu^{\prime \prime}(s) \mu(s)=0 \tag{17}
\end{equation*}
$$

Relations (15) and (17) imply second order differential equation

$$
\begin{equation*}
\mu^{\prime \prime}(s) \mu(s)+\mu^{\prime 2}(s)+1=0 \tag{18}
\end{equation*}
$$

whose general solution reads

$$
\begin{equation*}
\mu(s)=\sqrt{c^{2}-\left(s+c_{1}\right)^{2}} \tag{19}
\end{equation*}
$$

where $c \in R_{0}^{+}, c_{1} \in R$ and $|c|>\left|s+c_{1}\right|$. The arclength parameter of $\beta$ is given by

$$
t(s)=\int_{0}^{s}\left\|\beta^{\prime}(u)\right\| d u
$$

By using relations (13) and (19) we obtain

$$
\left\|\beta^{\prime}(s)\right\|=\frac{1}{\sqrt{1-\left(\frac{s+c_{1}}{c}\right)^{2}}}
$$

and therefore

$$
\begin{equation*}
t(s)=c \arcsin \left(\frac{s+c_{1}}{c}\right) \tag{20}
\end{equation*}
$$

Relations (14) and (20) imply

$$
\begin{equation*}
\beta(s)=\left(0, \cos \left(c \arcsin \left(\frac{s+c_{1}}{c}\right), \sin \left(c \arcsin \left(\frac{s+c_{1}}{c}\right)\right) .\right.\right. \tag{21}
\end{equation*}
$$

Substituting (19) and (21) in (13) and using reparametrization

$$
\begin{equation*}
u=\arcsin \left(\frac{s+c_{1}}{c}\right) \tag{22}
\end{equation*}
$$

we obtain that $\alpha$ is given by (11). In particular, by using (13),(19) and (21) we get

$$
\begin{equation*}
g\left(\alpha^{\prime \prime}(s), \alpha^{\prime \prime}(s)\right)=\frac{c^{2}\left[\left(1-c^{2}\right)\left(s+c_{1}^{2}\right)^{2}+c^{2}\left(c^{2}-1\right)\right]}{\left.c^{2}-\left(s+c_{1}\right)^{2}\right)^{3}} \tag{23}
\end{equation*}
$$

Since $g\left(\alpha^{\prime \prime}(s), \alpha^{\prime \prime}(s)\right) \neq 0$ it follows that $c \neq 1$ which proves statement (i).
(ii) Assume that $\alpha$ has null principal normal and that $\beta$ is a normal curve. As in the proof of statement (i) we easily obtain that $\alpha$ is given by (11). In particular, by using (23) it follows that $g\left(\alpha^{\prime \prime}, \alpha^{\prime \prime}\right)=0$ if $c=1$. Finally, putting $c=1$ in (11) we get that $\alpha$ is given by (12), which proves statement (ii).

Remark 1. By reparametrization $u=\arcsin \left(\frac{a}{c} \tan z\right), a^{2}=1-c^{2}, c \in R_{0}^{+}, c \neq 1$, parameter equation (11) can also be written as

$$
\alpha(z)=\frac{a}{\cos z} x(z),
$$

where

$$
x(z)=\frac{1}{a}\left(\sqrt{\cos ^{2} z-a^{2}}, \cos \left(c \arcsin \left(\frac{a}{c} \tan z\right)\right) \cos z, \sin \left(c \arcsin \left(\frac{a}{c} \tan z\right)\right) \cos z\right)
$$

is a unit speed spacelike curve lying in a pseudosphere $S_{1}^{2}(1)$.
Theorem 2. Let $\alpha$ be spacelike rectifying curve in $E_{1}^{3}$ and $\beta$ the orthogonal projection of $\alpha$ onto timelike plane of $E_{1}^{3}$.
(a) If $\alpha$ has non-null principal normal and $\beta$ is spacelike or timelike normal curve respectively, then up to a parametrization $\alpha$ is given by

$$
\alpha(u)=(\cosh (c u), \sinh (c u), c \cosh u)
$$

or by

$$
\alpha(u)=(\sinh (c u), \cosh (c u), c \sinh u),
$$

where $c \in R_{0}^{+}, c \neq 1$.
(b) If $\alpha$ has null principal normal and $\beta$ is spacelike or timelike normal curve respectively, then up to a parametrization $\alpha$ is given by

$$
\alpha(u)=(\cosh u, \sinh u, \cosh u)
$$

or by

$$
\alpha(u)=(\sinh u, \cosh u, \sinh u)
$$

(c) If $\beta$ is null normal curve, then $\alpha$ is the straight line.

Proof. In cases (a) and (b) the proof is similar to the proof of Theorem 1, so we omit it and give the proof only for the case (c). Assume that $\beta$ is the orthogonal projection of $\alpha$ onto timelike plane of $E_{1}^{3}$. Then $\alpha$ is given by

$$
\begin{equation*}
\alpha(s)=\beta(s)+\mu(s) \mathrm{v} \tag{24}
\end{equation*}
$$

where $s$ is arclength parameter of $\alpha, \mu(s)$ is some non-constant differentiable function and $\mathrm{v}=(0,0,1)$ is a unit spacelike vector. Assume that $\beta$ is a null normal curve. Then up to isometries of $E_{1}^{3}, \beta$ is the null straight line given by $\beta(t)=t(1,1,0)$. Differentiating relation (24) with respect to $s$ and using the condition $g\left(\alpha^{\prime}(s), \alpha^{\prime}(s)\right)=1$ we obtain $\mu(s)=s$. Hence $\alpha(s)=s(1,1,1)$ which means that $\alpha$ is the straight line.

In the next theorem we prove that the straight lines are the only rectifying curves in $E_{1}^{3}$ whose projection onto lightlike plane (with respect to a chosen screen distribution) is a normal curve.
Theorem 3. Let $\alpha$ be a spacelike rectifying curve with non-null principal normal in $E_{1}^{3}$ and $\beta$ the projection of $\alpha$ onto lightlike plane with the equation $x_{1}=x_{2}$. If $\beta$ is a normal curve and the screen distribution is given by (6) or (9), then $\alpha$ is the straight line.

Proof. Assume that $\beta$ is a normal curve and that the screen distribution is given by (6). By using relation (7) and the condition $g\left(\beta(s), \beta^{\prime}(s)\right)=0$ we find $\alpha_{3}(s) \alpha_{3}^{\prime}(s)=$ 0 . It follows that $\alpha$ is a planar curve and therefore the straight line.

Next, assume that the screen distribution is given by (9). By using relation (10) and the condition $g\left(\beta(s), \beta^{\prime}(s)\right)=0$ we obtain

$$
\left(\alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)\right)\left(\alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)\right)^{\prime}=0
$$

where $s$ is arclength parameter of $\alpha$. We distinguish two cases:
(1) $\alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)=0$;
(2) $\left(\alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)\right)^{\prime}=0$.
(1): $\alpha_{3}+\alpha_{1}-\alpha_{2}=0$. Then the curve $\alpha$ is given by

$$
\begin{equation*}
\alpha(s)=\left(\alpha_{1}(s), \alpha_{1}(s)+\alpha_{3}(s), \alpha_{3}(s)\right) \tag{25}
\end{equation*}
$$

so the condition $g\left(\alpha^{\prime}(s), \alpha^{\prime}(s)\right)=1$ implies

$$
\begin{equation*}
\alpha_{1}(s)=\frac{1}{2} \int \frac{d s}{\alpha_{3}^{\prime}(s)}-\alpha_{3}(s) \tag{26}
\end{equation*}
$$

From the relation $g\left(\alpha(s), \alpha^{\prime \prime}(s)\right)=0$ we find

$$
\begin{equation*}
\alpha_{1}^{\prime \prime} \alpha_{3}+\alpha_{1} \alpha_{3}^{\prime \prime}+2 \alpha_{3} \alpha_{3}^{\prime \prime}=0 \tag{27}
\end{equation*}
$$

Relations (26) and (27) imply

$$
\alpha_{3}^{\prime \prime}(s)\left(-\frac{\alpha_{3}(s)}{\alpha_{3}^{\prime 2}(s)}+\int \frac{d s}{\alpha_{3}^{\prime}(s)}\right)=0
$$

We distinguish two subcases:
(1.1) $\alpha_{3}^{\prime \prime}(s)=0 ;$
(1.2) $-\frac{\alpha_{3}(s)}{\alpha_{3}^{\prime 2}(s)}+\int \frac{d s}{\alpha_{3}^{\prime}(s)}=0$.
(1.1): If $\alpha_{3}^{\prime \prime}(s)=0$, by using (25) and (26) we obtain that $\alpha$ has parameter equation

$$
\alpha(s)=\left(\frac{s-2 a^{2} s-2 a b}{2 a}, \frac{s}{2 a}, a s+b\right), \quad a \in R_{0}, b \in R
$$

and thus it is the straight line.
(1.2): If

$$
-\frac{\alpha_{3}(s)}{\alpha_{3}^{\prime 2}(s)}+\int \frac{d s}{\alpha_{3}^{\prime}(s)}=0
$$

differentiating the last equation with respect to $s$ we find $\alpha_{3}(s) \alpha_{3}^{\prime \prime}(s)=0$. Consequently, this subcase reduces to the subcase (1.1), which means that $\alpha$ is the straight line.
(2): $\left(\alpha_{3}(s)+\alpha_{1}(s)-\alpha_{2}(s)\right)^{\prime}=0$. Then $\alpha$ is given by

$$
\alpha(s)=\left(\alpha_{1}(s), \alpha_{1}(s)+\alpha_{3}(s)-c, \alpha_{3}(s)\right), \quad c \in R,
$$

and therefore it is congruent to the curve given by (25). Hence $\alpha$ is the straight line, which proves the theorem.

Remark 2. The projected curve $\beta$ in Theorem 3 is the null straight line.

## 4. Some relations between non-planar spacelike normal curves and planar rectifying curves in $\mathbb{E}_{1}^{3}$

It is known that every non-planar spacelike normal curve in $E_{1}^{3}$ lies in some quadric ([5]). In this section, we obtain some explicit parameter equations of non-planar spacelike normal curves in $E_{1}^{3}$ assuming that their projection onto lightlike plane of $E_{1}^{3}$, with respect to a chosen screen distribution, is a rectifying W-curve. In Euclidean 3 -space, it is known that for rectifying curves the curvature ratio $\kappa_{2}(s) / \kappa_{1}(s)$ is a non-constant linear function ([1]). The same property hold for timelike and spacelike rectifying curves with non-null principal normal in Minkowski 3-space ([7]). On the other hand, rectifying curves with null principal normal in $E_{1}^{3}$ are planar curves lying fully in the lightlike plane $\{T, N\}$. Therefore, every rectifying curve with null principal normal in $E_{1}^{3}$ is a pseudo null curve. The following theorem is given in [11] for pseudo null W-curves.

Theorem 4 ([11, Theorem A]). All pseudo null spacelike curves in $E_{1}^{3}$ with constant curvatures can be classified as:
(1) $\kappa_{1}(s)=0$ if and only if $\beta$ is a part of a spacelike straight line;
(2) $\kappa_{1}(s)=1$ and $\kappa_{2}(s)=0$ if and only if $\beta$ is a part of a planar curve with parametrization

$$
\begin{equation*}
\beta(s)=\left(\frac{s^{2}}{2}, \frac{s^{2}}{2}, s\right) ; \tag{28}
\end{equation*}
$$

(3) $\kappa_{1}(s)=1$ and $\kappa_{2}(s)=c_{2}=$ constant $\neq 0$ if and only if $\beta$ is a part of a planar curve with parametrization

$$
\begin{equation*}
\beta(s)=\frac{1}{c_{2}^{2}}\left(\cosh \left(c_{2} s\right)+\sinh \left(c_{2} s\right), \cosh \left(c_{2} s\right)+\sinh \left(c_{2} s\right), c_{2}^{2} s\right) \tag{29}
\end{equation*}
$$

In the next theorem, we obtain explicit parameter equation of non-planar spacelike normal curve in $E_{1}^{3}$ whose projection onto lightlike plane of $E_{1}^{3}$ (with respect to a chosen screen distribution) is a rectifying W-curve.

Theorem 5. Let $\alpha$ be a non-planar spacelike normal curve in $E_{1}^{3}$ and $\beta$ the projection of $\alpha$ onto lightlike plane with the equation $x_{1}=x_{2}$ and the screen distribution given by (6). If $\beta$ is a rectifying $W$-curve with parameter equation (28) or (29) respectively, then $\alpha$ is given by

$$
\begin{equation*}
\alpha(s)=\left(-\frac{c}{s^{2}}+\frac{s^{2}}{2}+\frac{1}{2}, \frac{c}{s^{2}}+\frac{s^{2}}{2}-\frac{1}{2}, s\right), \quad s^{2}>8 c, \quad c \in R, \tag{30}
\end{equation*}
$$

or else by

$$
\alpha(s)=\left(\frac{2-c_{2}}{c_{2}^{2}} e^{c_{2} s}-c e^{-c_{2} s}+\frac{c_{2}^{2}}{4} s^{2} e^{-c_{2} s}, \frac{1}{c_{2}} e^{c_{2} s}+c e^{-c_{2} s}-\frac{c_{2}^{2}}{4} s^{2} e^{-c_{2} s}, s\right)
$$

where $c_{2} \in R_{0}, c \in R$.
Proof. First assume that $\beta$ is given by (28). From relations (7) and (28) we obtain that $\alpha$ is given by

$$
\begin{equation*}
\alpha(s)=\left(s^{2}-\alpha_{2}(s), \alpha_{2}(s), s\right) \tag{31}
\end{equation*}
$$

where $\alpha_{2}(s)$ is some differentiable function and $s$ is arclength parameter of $\beta$. By using the condition $g\left(\alpha(s), \alpha^{\prime}(s)\right)=0$ we obtain linear differential equation

$$
\alpha_{2}^{\prime}(s)+\frac{2}{s} \alpha_{2}(s)=2 s-\frac{1}{s}, \quad s \neq 0
$$

whose general solution reads

$$
\alpha_{2}(s)=\frac{c}{s^{2}}+\frac{s^{2}}{2}-\frac{1}{2}, \quad c \in R
$$

Substituting this in (31) we find that $\alpha$ has parameter equation (30).
If $\beta$ is given by (29), the proof is analogous.
Remark 3. The curve given by (30) lies in pseudosphere $-x^{2}+y^{2}+z^{2}=2 c$ if $c>0$, lightcone $-x^{2}+y^{2}+z^{2}=0$ if $c=0$ and pseudohyperbolic space $-x^{2}+y^{2}+z^{2}=2 c$ if $c<0$.

The last theorem can be proved in a similar way.

Theorem 6. Let $\alpha$ be a non-planar spacelike normal curve in $E_{1}^{3}$ and $\beta$ the projection of $\alpha$ onto lightlike plane with the equation $x_{1}=x_{2}$ and the screen distribution given by (9). If $\beta$ is a rectifying $W$-curve with parameter equation (28) or (29) respectively, then $\alpha$ is given by

$$
\alpha(s)=\left(\frac{1}{s\left(1+\frac{s}{2}\right)}\left(c+\frac{s^{4}}{4}+\frac{s^{3}}{2}+\frac{s^{2}}{2}\right), \frac{s^{2}}{2}, s\left(1+\frac{s}{2}\right)-\frac{1}{s\left(1+\frac{s}{2}\right)}\left(c+\frac{s^{4}}{4}+\frac{s^{3}}{2}+\frac{s^{2}}{2}\right)\right)
$$

where $c \in R$, or else by

$$
\begin{aligned}
\alpha(s)= & \left(\frac{1}{s+\frac{1}{c_{2}} e^{c_{2} s}}\left(c+\frac{1}{c_{2}^{2}} e^{2 c_{2} s}+\frac{s^{2}}{2}+\frac{s}{c_{2}} e^{c_{2} s}\right), \frac{1}{c_{2}} e^{c_{2} s}, s+\frac{1}{c_{2}} e^{c_{2} s}\right. \\
& \left.-\frac{1}{s+\frac{1}{c_{2}} e^{c_{2} s}}\left(c+\frac{1}{c_{2}^{2}} e^{2 c_{2} s}+\frac{s^{2}}{2}+\frac{s}{c_{2}} e^{c_{2} s}\right)\right), \quad c_{2} \in R_{0}, c \in R .
\end{aligned}
$$

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[^0]:    *Corresponding author. Email addresses: grbovic@yahoo.com (M. Grbović), emilija@kg.ac.rs (E. Nešović)

