

## Some low energy restrictions for a flipped left-right symmetric $E_6$ model

J.O. Eeg

Dept. of Physics, University of Oslo,  
P.O.Box 1048 Blindern, N-0316 Oslo 3, Norway

### Abstract

I give a short review of a new left-right symmetric  $E_6$  model containing a right-handed  $W$ -boson  $W_R$  with odd R-parity. Some restrictions from low-energy data are presented, the most important: coming from  $D - \bar{D}$  mixing. It is found that  $W_R$  is likely to be heavier than 500 to 1500 GeV.

The Standard Model (SM) gives for the moment very good agreement with data. However, elementary particle physicists hardly think that the SM is the final answer, and are eager to see if there are effects due to new physics beyond the SM. Such effects might occur as signals in collisions at very high energy, or as very exotic decay processes. A possible extension of the SM is the left-right symmetric model containing an extra gauge-boson  $W_R$  mediating charged right-handed currents[1]. Such interactions have significant impact on low-energy data, in particular  $K - \bar{K}$  mixing[2], and it was found that the mass  $M_R$  of  $W_R$  must be bigger than 1-10 TeV to avoid conflicts with data. Thus the prospects to observe such a particle in a relatively near future is rather poor.

Left-right symmetric models may also be embedded in GUT-models, like  $SO(10)$  and  $E_6$ . Due to the heterotic string[3],  $E_6$  models have recently been extensively studied. Low energy restrictions for an  $E_6$  model broken down to  $SU(3) \times SU(2) \times U(1) \times U(1)$  have been studied[4] previously. But  $E_6$  may also be broken down to  $G = SU(3) \times SU(2) \times SU(2) \times U(1)$ , containing a left-right symmetric model. A conventional particle

assignment for one generation of fermions would then be

$$\begin{aligned}
 \begin{pmatrix} u \\ d \end{pmatrix}_L &: (3, 2, 1, \frac{1}{6}), & (d^c, u^c)_L &: (\bar{3}, 1, 2, -\frac{1}{6}), \\
 h_L &: (3, 1, 1, -\frac{1}{3}), & h_L^c &: (\bar{3}, 1, 1, \frac{1}{3}), \\
 \begin{pmatrix} \nu_e \\ e \end{pmatrix}_L &: (1, 2, 1, -\frac{1}{2}), & (e^c, N^c)_L &: (1, 1, 2, \frac{1}{2}), \\
 \begin{pmatrix} \nu_E & E^c \\ E & N_E^c \end{pmatrix}_L &: (1, 2, 2, 0), & n_L &: (1, 1, 1, 0),
 \end{aligned} \tag{1}$$

where the superscript  $c$  denotes charge conjugation. With this particle assignment one obtains the ordinary left-right symmetric model with some extension due to the exotic fermions contained in the  $2\bar{1}$  representation of  $E_6$ .  $h$  is an exotic quark with electric charge  $-1/3$ , and  $E, \nu_E, N_E, n$  are exotic non-hadronic fermions (-hereafter called "leptons");  $E$  has electric charge  $-1$ , while the rest are neutral.

Recently, Ma[5] proposed to flip the particle content of (1) in the following way:

$$\begin{aligned}
 \begin{pmatrix} u \\ d \end{pmatrix}_L &: (3, 2, 1, \frac{1}{6}), & d_L^c &: (\bar{3}, 1, 1, \frac{1}{3}), \\
 h_L &: (3, 1, 1, -\frac{1}{3}), & (h_L^c, u^c)_L &: (\bar{3}, 1, 2, -\frac{1}{6}), \\
 \begin{pmatrix} \nu_e & E^c \\ e & N_E^c \end{pmatrix}_L &: (1, 2, 2, 0), & (e^c, n)_L &: (1, 1, 2, \frac{1}{2}), \\
 \begin{pmatrix} \nu_E \\ E \end{pmatrix}_L &: (1, 2, 1, -\frac{1}{2}), & N_L^c &: (1, 1, 1, 0).
 \end{aligned} \tag{2}$$

Note that this flipping is not done at the level of unbroken  $E_6$ , but at the level of  $G$ . In this sense the Ma-model[5] is different from the flipped  $SU(5) \times U(1)$  model[6]. The particle assignment (2) means that due to  $W_R$  exchange,

$$u \leftrightarrow h, \nu_e \leftrightarrow E^c, e \leftrightarrow N_E^c, e^c \leftrightarrow n, \tag{3}$$

while there is *no*  $W_R$ -coupling for  $u \leftrightarrow d$  as in ordinary left-right symmetric models. One should note that the exotic "leptons" have lepton numbers equal to zero, while  $W_R$  and  $h$  have non-zero lepton numbers.

The SM particles have even R-parity, and their SUSY-partners odd R-parity, as usual. The exotic quark  $h$ , the exotic "leptons"  $E, \nu_E, N_E, n, N$  and the righthanded W-boson  $W_R$  all have odd R-parity, and consequently their SUSY-partners have even R-parity. These R-parity assignments implies that the SUSY-partners of the exotic neutral "leptons" may acquire v.e.v.'s and be used as Higgses. An important consequence of the odd R-parity of  $W_R$  is that there is no mixing between  $W_R$  and  $W_L$  as in ordinary left-right symmetric models. This property combined with the absence of a  $duW_R$  (and  $suW_R$ ) coupling means that there is no significant contribution from  $W_R$ -exchange to  $K - \bar{K}$  mixing. Thus the previous bound[2] on  $W_R$  from  $K - \bar{K}$  mixing does not apply any more, and  $W_R$  could a priori be as light as 300 GeV, say[5,7,8,9,10].

As in most models beyond the SM, in particular in models including SUSY, rare effects will be generated also in the Ma model[5]. Such effects will depend on the details

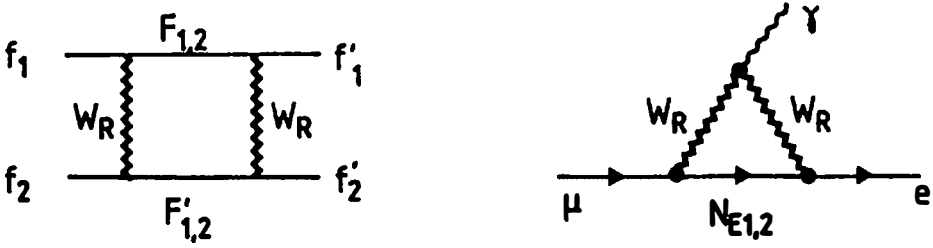


Figure 1: Fig.1.a(left): The box diagram. The exotic fermions  $F$  and  $F'$ , and  $W_R$  are running in the loop, while the external fermions  $f_i$  ( $i = 1, 2, 1', 2'$ ) are SM fermions. Fig.1b(right): Diagram for  $\mu \rightarrow e\gamma$  due to  $W_R$ -exchange.

of the model[11,12,13], in particular the Yukawa interactions, which may be chosen in slightly different ways. One should note that one may also construct non-SUSY versions of the model[14]. I have therefore concentrated on effects due to exchanges of  $W_R$ , which are expected[10] to set the typical scale of rare effects in the model.

Besides the limits on the masses of the exotic particles which might be obtained in high energy collisions[7,8,9], there will also be low energy limits obtained from absence of rare effects. Several rare effects, such as  $D - \bar{D}$  mixing,  $D^0 \rightarrow \mu\bar{\mu}$ ,  $D^0 \rightarrow \mu\bar{e}$ ,  $\mu \rightarrow 3e$ , and  $\mu N \rightarrow eN$ , are obtained [15] from box-diagrams with double  $W_R$ -exchange, as shown in Fig.1a. The most important case turns out to be  $D - \bar{D}$  mixing[10,15].

For the  $D - \bar{D}$  mixing case  $f_1 = c$ ,  $f'_1 = u$ ,  $f_2 = \bar{u}$ ,  $f'_2 = \bar{c}$ , in Fig.1a. The fermions in the loop are the  $h$ -quarks of three different generations. The mixing among these are analogous to the mixing of the  $d, s, b$  quarks of the SM. Neglecting CP-violating effects, it suffices to consider two generations only, and the diagram is analogous to that for  $K - \bar{K}$  mixing with the substitutions  $s \rightarrow c$ ;  $d \rightarrow u$ ;  $u, c \rightarrow h_{1,2}$ ;  $W_L \rightarrow W_R$ . (Note that exchange of one  $W_L$  and one  $W_R$  is forbidden.) Recall that in the SM,  $D - \bar{D}$  mixing is very small, and moreover hard to estimate due to non-perturbative effects. Comparing the box diagram for  $D - \bar{D}$  mixing due to  $W_R$ -exchange (R) with the short distance (SD) part of the corresponding box diagram for  $K - \bar{K}$  mixing in the SM, we obtain

$$\frac{(\Delta m_D)_R}{(\Delta m_K)_{SD}} = \frac{m_D}{m_K} \left( \frac{f_D}{f_K} \right)^2 \frac{B_D}{B_K} \frac{\eta_D}{\eta_K} \left( \frac{\lambda_R}{\lambda_L} \right)^2 \frac{Q_h}{m_c^2} \frac{(g_R M_L)^4}{(g_L M_R)^4}, \quad (4)$$

where  $\Delta m_K = m_{K_L} - m_{K_S}$ , and similarly for  $D^0$ .  $M_{L,R}$  are the masses and  $g_{L,R}$  the gauge couplings of  $W_{L,R}$ .  $\lambda_{L,R}$  are the Cabibbo factors for  $h_{1,2}$  mixing and  $d, s$  mixing, respectively. (One may for instance assume  $\lambda_R \simeq \lambda_L$  [10], or use the Fritzsch ansatz[16]).  $B_K$  is the so-called  $B$ -parameter due to the hadronic matrix element, and  $\eta_K$  contains the perturbative QCD corrections in  $K - \bar{K}$  mixing[17].  $B_D$  and  $\eta_D$  are the corresponding parameters for  $D - \bar{D}$  mixing in our case.  $Q_h$  is a factor determined by the box loop integration:

$$Q_h \simeq M_2^2; \quad Q_h \simeq \frac{1}{3} M_R^2; \quad Q_h \simeq M_R^2, \quad (5)$$

for the cases  $M_1^2 \ll M_2^2 \ll M_R^2$ ,  $M_1^2 \ll M_2^2 \sim M_R^2$ , and  $M_2^2 \gg M_{1,R}^2$  respectively. ( $M_{1,2}$  are the masses of  $h_{1,2}$ .) It is known[17] that the SD amplitude for  $K^0 \rightarrow \bar{K}^0$  can account for  $\simeq 1/2$  of the observed  $\Delta m_K$ . Thus,  $\Delta m_K = \beta(\Delta m_K)_{SD}$  with  $\beta \simeq 2$ .

Using[18]  $\Delta m_D/\Delta m_K < 37$ , we obtain a bound on the combination of parameters of the right hand side of (4). Taking explicit values for  $M_{1,2}$ , we will obtain a corresponding bound for the  $W_R$  mass. In [7,8,9]  $M_1 \sim 100$  and  $300$  GeV are taken as examples. One would then obtain  $M_R > 0.5$  to  $1.5$  TeV. (For numerical estimates,  $g_R = g_L$  is assumed. The ratios of the  $f$ 's,  $B$ 's and  $\eta$ 's in (4) are of order one[10,15]). Thus, unless the masses  $M_1$  and  $M_2$  are rather close,  $W_R$  has to be heavier than the bound 2-300GeV mentioned in [5,7,8,9,10]. (If  $M_1/M_2 \simeq 0.7$ , the  $W_R$ -mass may still be 2-300 GeV [10], but it might be argued that  $M_2^2 \gg M_1^2$  is more likely, and more specific that  $M_1 : M_2 : M_3 = m_d : m_s : m_b$  at some scale[14])

$W_R$ -exchange will also contribute to  $\mu \rightarrow e\gamma$ , as shown in Fig.1b, but this process gives a slightly milder bound on  $M_R$  (350 to 650 GeV, say) compared to (4)[15]. (Note that in this case there is a complication due to mixing between  $N_E^c, \nu_E$  and  $n$  of the same generation.) The process  $\mu \rightarrow 3e$  gives a bound comparable to  $\mu \rightarrow e\gamma$ , while  $\mu N \rightarrow eN$  gives an uninteresting bound  $\sim 200$  GeV for  $M_R$  ( $M_2^2 \gg M_1^2$  is assumed). The process  $D^0 \rightarrow \mu\bar{\mu}$  will not give an interesting bound on  $M_R$ . One finds that for  $M_R \sim 0.5$  to  $1$  TeV,  $Br(D^0 \rightarrow \mu\bar{\mu})$  will hardly exceed  $10^{-10}$  (-but will still be bigger than in the SM), while  $Br(D^0 \rightarrow \mu\bar{e})$  is one or two orders of magnitude smaller.  $W_R$ -exchange will also generate magnetic (transition) moments of the neutrinos of the SM. But as in the SM, the moments are proportional to the neutrino mass, and will be too small to be interesting for the proposed VVO-solution of the solar neutrino problem[19].

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