

MEAN SPHERICAL APPROXIMATION FOR A SYSTEM OF PARTICLES
INTERACTING WITH DAMPED OSCILLATORY POTENTIAL

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An important characteristic of the interaction between the ions in a liquid metal is the presence of oscillations in the potential energy as a function of the interparticle distance /1/. In order to calculate the structure factor, different model potentials were proposed and by comparing them, one can notice a considerable distinction between them /2/. It is a characteristic feature of all the calculations performed so far that the interaction potential is cut-off at certain distance and no more than one oscillation is actually taken into account.

Although it has a number of deficiencies /3/ the so-called mean spherical approximation /MSA/ represents a frequently used approximation for different model potentials. The main argument in favour of the MSA is that it is exactly solvable in many cases, best known example being the Percus-Yevick equation for hard-core particles.

In the following we consider a system of particles interacting with the potential

$$V_i(r) = \begin{cases} A e^{-\kappa r} \frac{\sin(lr+m)}{r} & r < R \\ 0 & r > R \end{cases} \quad /a/$$

where $A, \kappa, l,$ and m are parameters, r is the interparticle distance and R is the radius of the hard core of the particles. The above potential differs from the Friedel potential. The oscillations are damped by a factor $\frac{e^{-\kappa r}}{r}$ instead of $\frac{1}{r^2}$. The presence of an exponential damping may be expected because the Fermi surface in liquid metals is not as sharp as in solid metals /1/.

Accordingly we start from the equation of Ornstein and Zernike

$$h(r) = c(r) + \rho \int h(|\vec{r}-\vec{r}'|) c(|\vec{r}'|) d\vec{r}' \quad /b/$$

where $h(r) = g(r) - 1$ and $g(r)$ is the pair distribution function, $c(r)$ is the direct correlation function and ρ is the number

particle density. The equation /b/ is closed by the impenetrability condition $g(r) = 0$ for $r < R$ and by the only assumption of MSA that for $r > R$ holds $c(r) = -\beta V(r)$ where β is the inverse temperature $\beta = \frac{1}{kT}$.

Following the standard procedure the initial problem is reduced to a system of six algebraic equations, which will be given elsewhere.

The following expression for the direct correlation function for $r < R$ is found

$$\chi c_o(x) = v_0 + v_1 x + v_2 x^2 + v_3 x^4 + v_4 e^{\kappa x} \cos \lambda x + v_5 e^{\kappa x} \sin \lambda x + v_6 e^{-\kappa x} \cos \lambda x + v_7 e^{-\kappa x} \sin \lambda x / c,$$
 where $\kappa = \frac{z}{R}$, $\lambda = l \cdot R$, $\kappa = \kappa R$, $c_o(r) = c(r) + \beta V(r)$ and the coefficients v_i are given through the solution of the algebraic equations. These equations were solved numerically for some particular values of the parameters using the gradient and the Newton-Raphson methods.

Taking Fourier transform of /b/ and using the result /c/ and the MSA assumption it is easy to compute the structure factor. Comparing our results with the structure factor of Ashcroft and Lekner /4/ we conclude that the positions of the maxima and minima are unchanged. The amplitudes are changed only for low q , $q = k \cdot R$ and k is the wave number, the reduction of the height of the first maximum being most pronounced. The behaviour for large q is dominated by the hard-core repulsion and is almost unaffected by the presence of /a/. Presumably better agreement with experiment can be obtained if some soft repulsion is added and this can be tackled within MSA if we assume potential of the form

$$V = V_1 + B \frac{e^{-\eta r}}{r}, \quad B > 0$$

References

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